

Latest technical developments at THEMIS of possible interest for EST

B. Gelly – CNRS



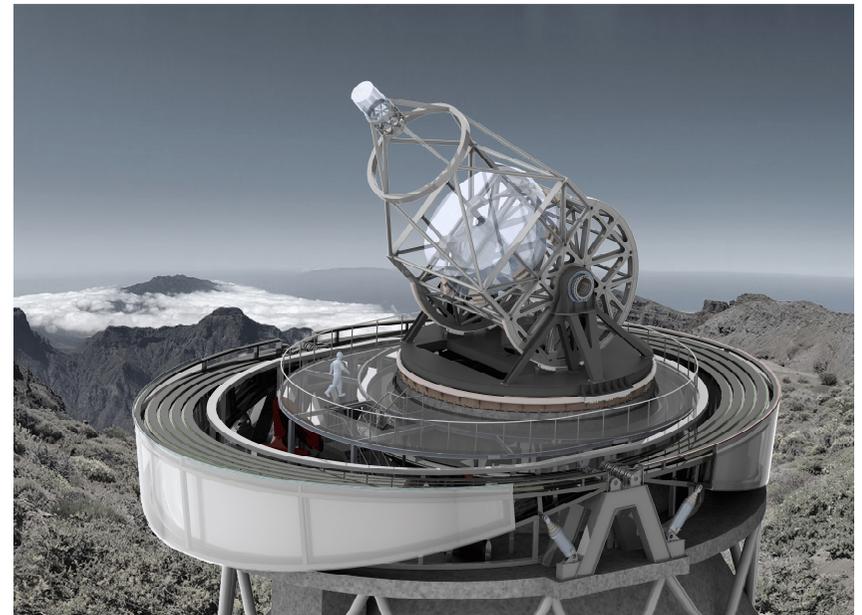
International Research Lab 2009

FSLAC

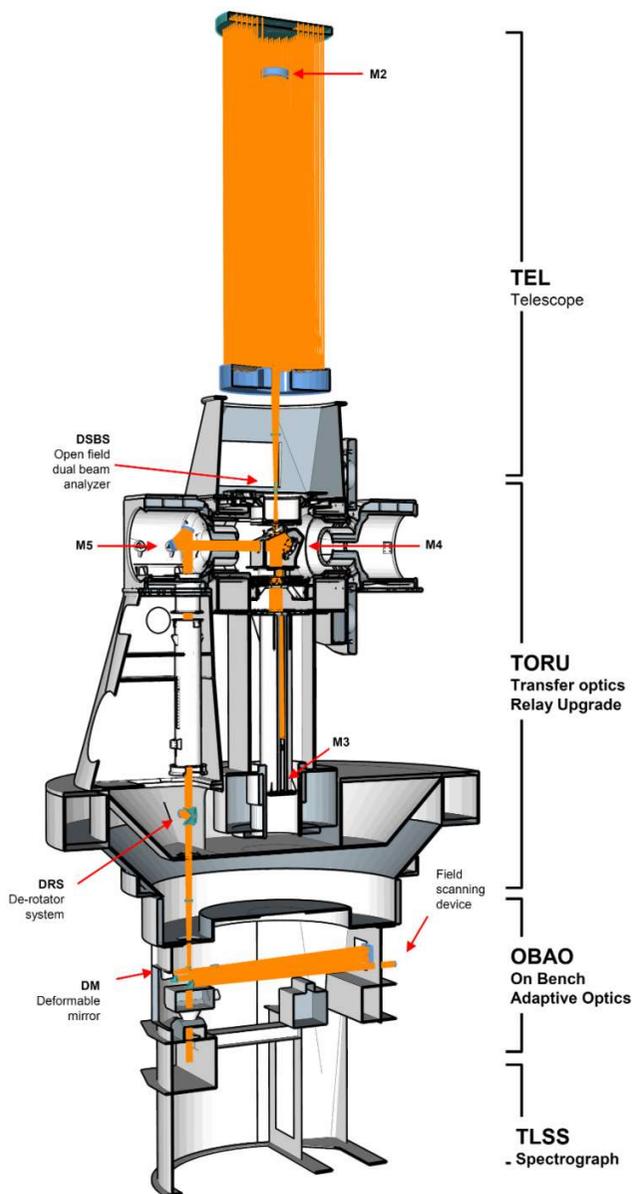
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THEMIS overhaul: 2016 → 2020



- THEMIS has been widely renewed since 2016 and over 5 years, resulting in a **big success for French teams**, running on EU money (~1M€ funding came from two SOLARNET EU programs)
- Some of the technical developments we have made are of possible interest for EST
- **Superimposed dual-beam polarimetric analysis without field limitation (1' polarimetric field) (Semel M., Lopez A., Le Men C. & THEMIS)**
- **“Polarization- friendly” complete redesign of the whole transfer optics (Le Men, C. & THEMIS)**
- **TAO “classical” (single-DM) adaptive optics based on innovative wavefront sensing and mirror commanding concepts that can be applied to larger systems (The AIRI team @CRAL & THEMIS)**
- Instrumental and strategic know-how is as good as ever was
- Teams to build and operate do exist
- Bosses to pay some small running costs also can be found

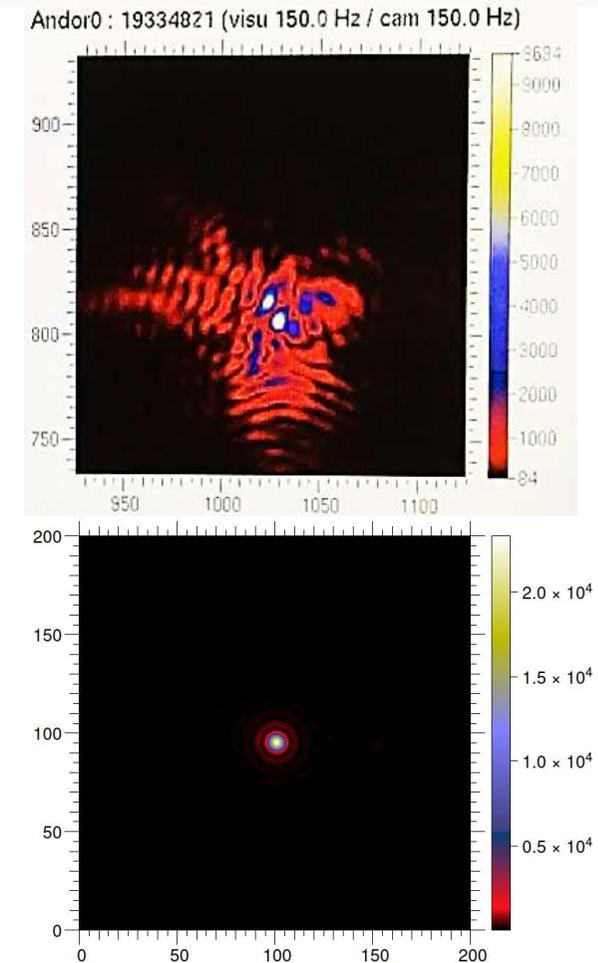
Themis Adaptive Optics TAO initial specifications and objectives

- Specifications

- 90 cm diameter solar telescope;
- **76 sub-aperture Shack-Hartmann wavefront sensor** (10×10);
- 380×380 pixel WFS images, Mikrotron EoSens 4CXP detector;
- **97 actuator Alpao DM** (11×11);
- RTC computer: CPU i7-4790K (Q2'14) at 4.2 GHz, 4 cores, up to 50 Gflops/core **with AVX2 + FMA instructions.**

- Objectives

- **close the AO loop on the Sun** (started from scratch mid-2016, job done December 2020);
- **unsupervised AO system** (optimal correction whatever the conditions)
- flexible RTC software to implement and experiment new algorithms (Julia lang)
- **RTC software running in CPU (no GPU) @1250 Hz**



Thiebaut, E., Tallon, M. et al, SPIE proceedings 2022

TAO linearized model of wavefront sensor data and DM commands in the THEMIS system

WFS data Actual wavefront Actuator commands ... some noise

$$\mathbf{d}_t = \mathbf{S} \cdot (\mathbf{w}_t + \mathbf{M} \cdot \mathbf{a}_t) + \mathbf{z}_t$$

Sensor linear response Mirror influence matrix

$\mathbf{G} = \mathbf{S} \cdot \mathbf{M}$ is the **interaction matrix**

$\mathbf{y}_t = \mathbf{d}_t - \mathbf{G} \cdot \mathbf{a}_t$ are \approx open loop data

$\mu_t > 0$ and $\mathbf{W} \approx \text{Cov}(w_t)$ are **loop parameters** to enforce **spatial regularization**

$\rho_t > 0$ is a **loop parameter** to impose **temporal continuity**

- Wavefront is represented in the basis of influence functions of the DM
- No predictions in TAO v.2 (but will appear in TAO v.3)

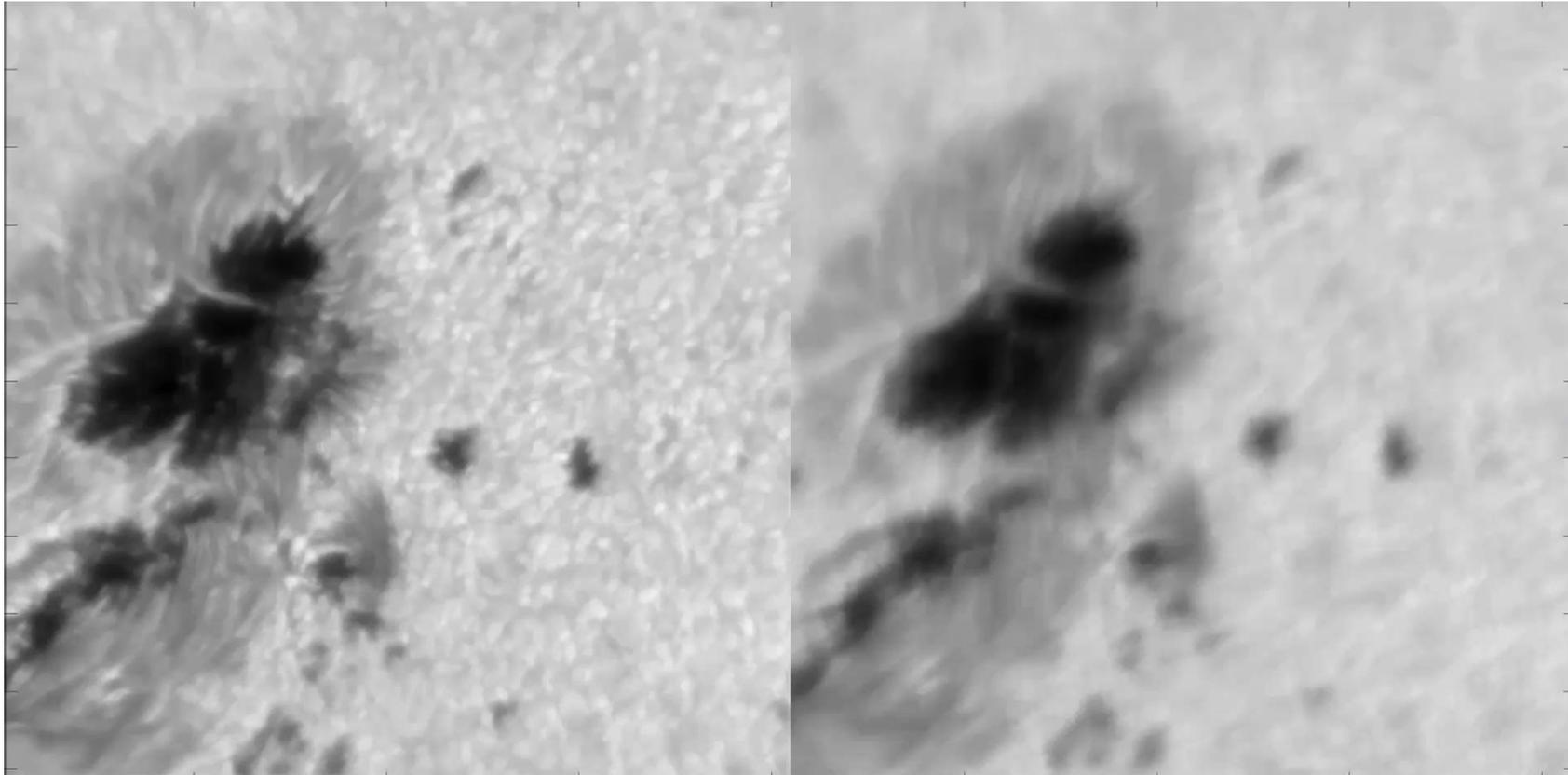
$$\mathbf{a}_{t+\delta t} = \underset{\text{argmin}}{\{ \|\mathbf{y}_t + \mathbf{G} \cdot \mathbf{a}\|_{\text{Cov}(\mathbf{z}_t)^{-1}}^2 + \mu_t \|\mathbf{a}\|_{\mathbf{W}}^2 + \rho_t \|\mathbf{a} - \mathbf{a}_t\|^2 \}}$$

→ Requires to solve an inverse problem (*argmin...*) at each step (all terms may change with time)

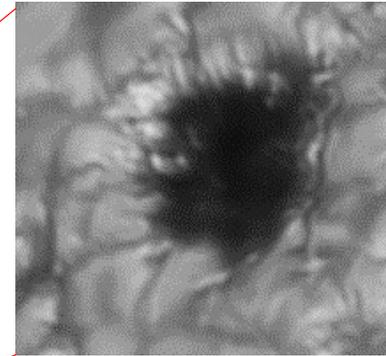
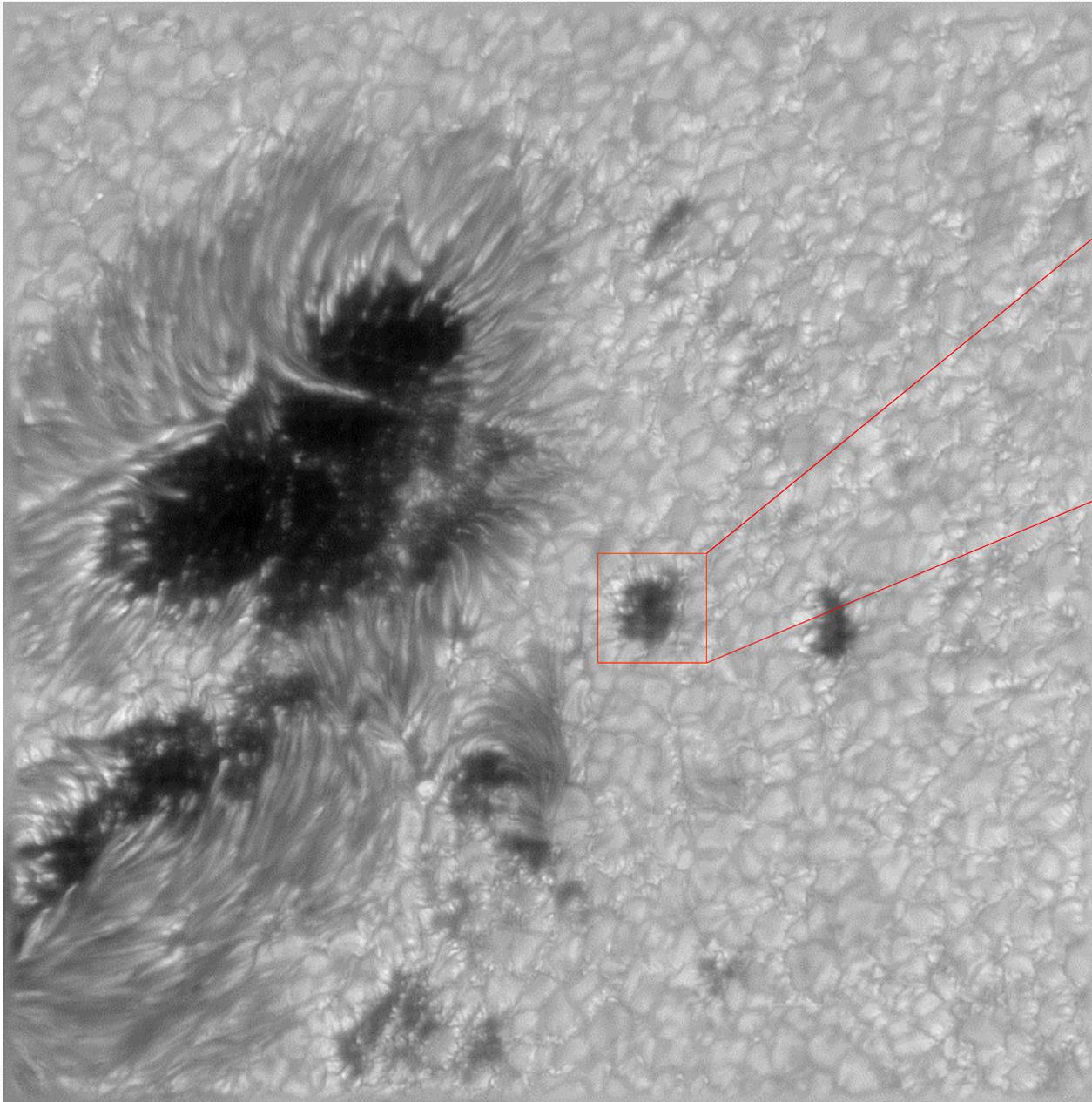
Thiebaut, E., Tallon, M. et al, SPIE proceedings 2022

TAO going live on NOAA 12975

- Successive recordings in equivalent seeing conditions
- 55'' square field, 20 fps (0.3 ms), 2k x 2k (0.03'' /px)



TAO + TSI : NOAA 12975



- 31/03/ 2022
- with post processing THEMIS Knox-Thompson (TSI)
- 55X55" FOV
- **0.17 " res (0.035 " pixel) near diffraction limit**
- 630 / 1 nm red filter

TAO v2.5 on granulation

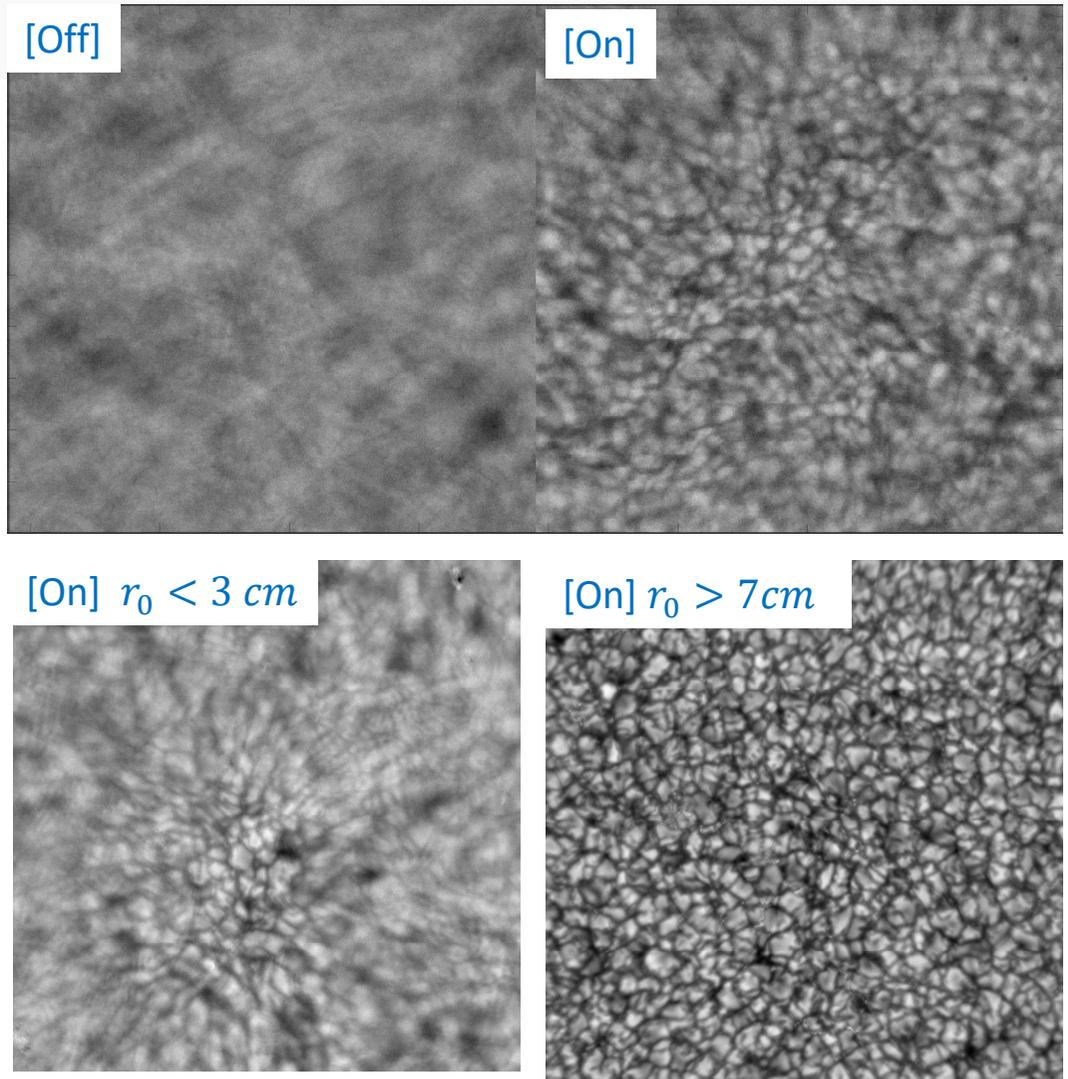
- Isoplanatic patch can be severely reduced by high altitude turbulence
- Ground layer conjugated AO show limitations coming from high altitude seeing not detected by

Given:

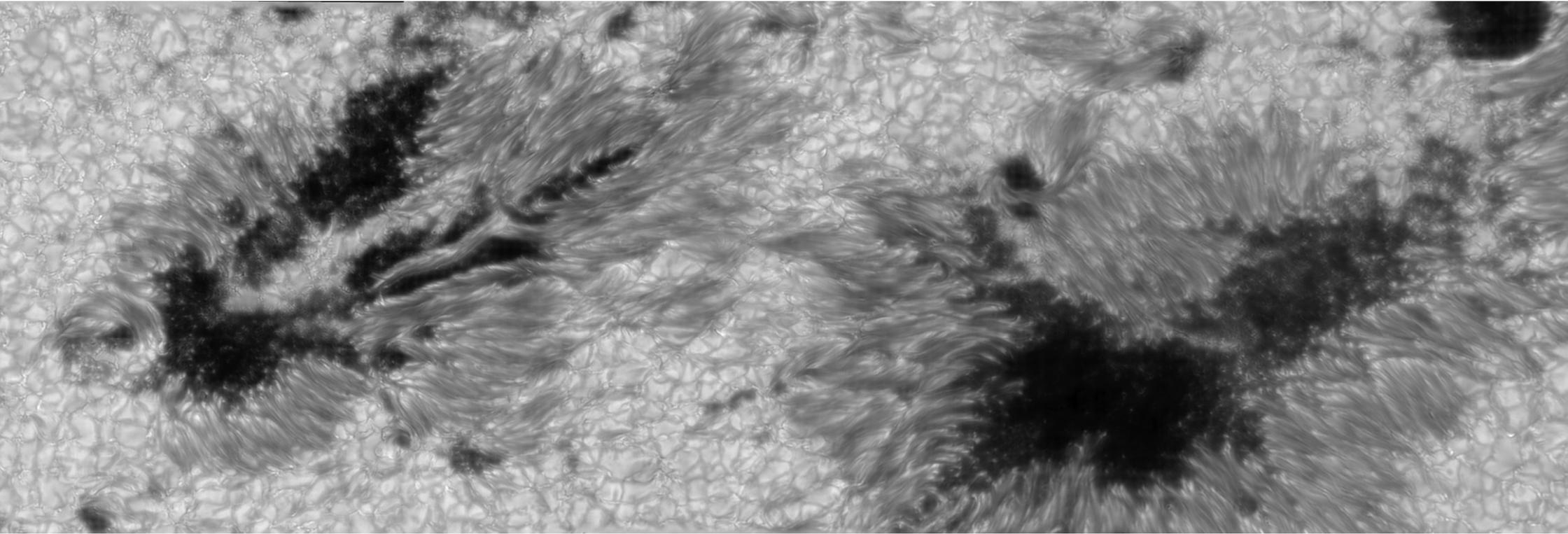
$$\alpha = r_0 / D$$

the same kind of result on a 4m telescope would require $r_0 = 28 \text{ cm}$

~ 55"



TAO + TSI NOAA 13354 wide field

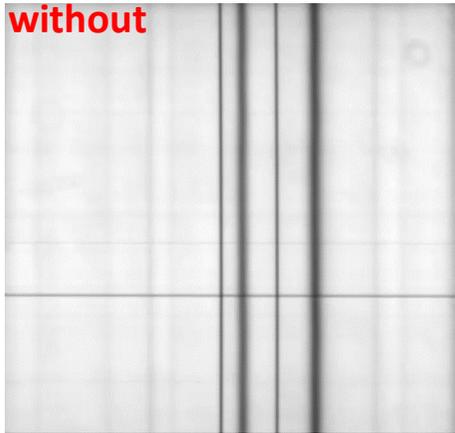


- NOAA 13354 from 2023/06/28
- 180x50 " field
- **0.17 " res (0.035 " pixel) near diffraction limit**
- 630 / 1 nm red filter
- Solar telescope **with robust and innovative AO**
- Record bursts **and restore from residual seeing**
- Stitch
- => Routine operation with pipeline !

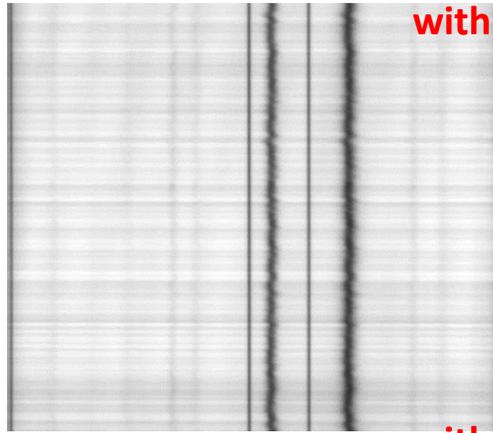
TAO for spectropolarimetry

110" spatial field (0.06"/px), ~5.5 A spectral field, 3mA /px

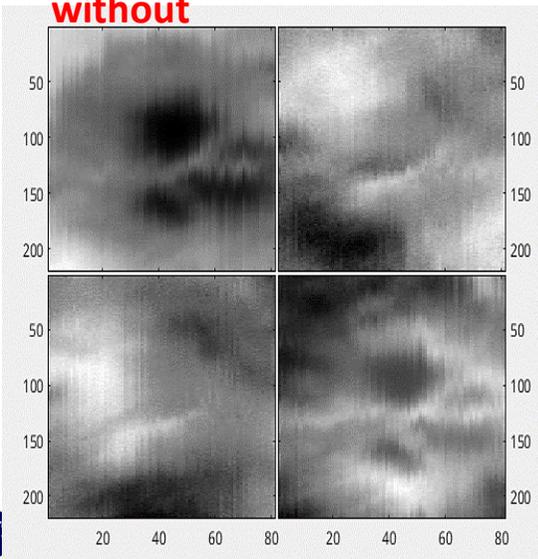
without



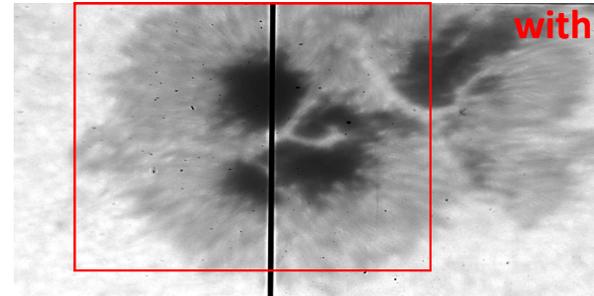
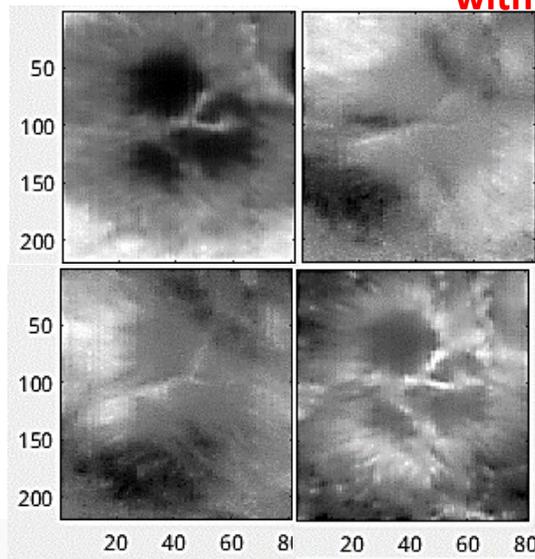
with



without



with

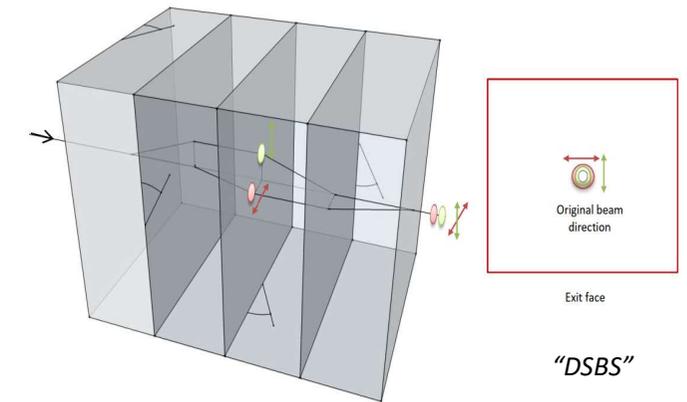
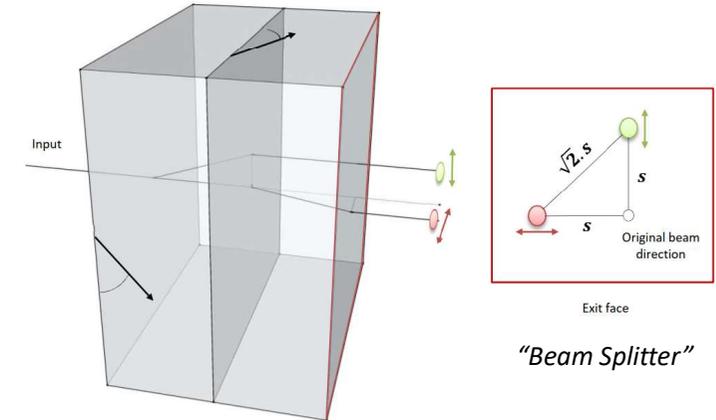
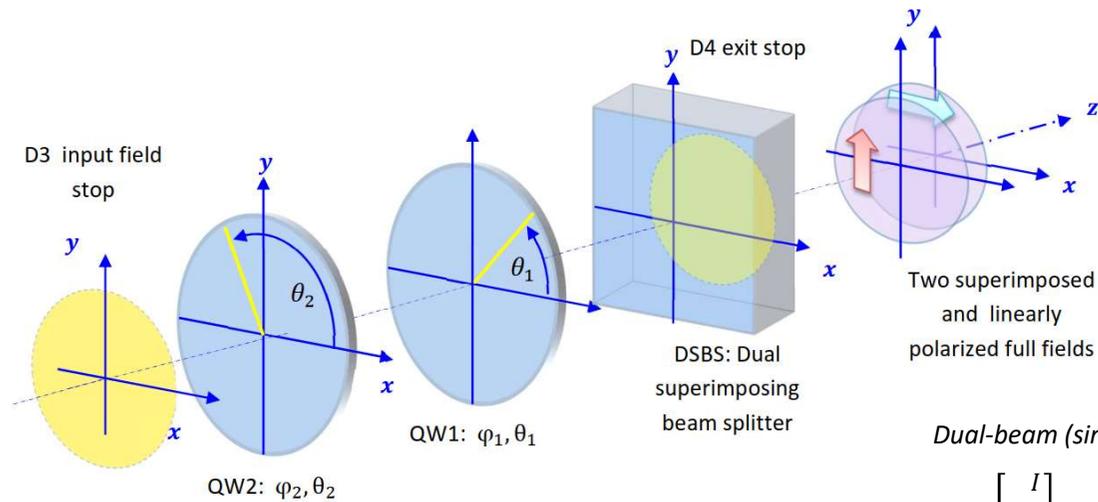


- **Stokes maps** are the last necessary step before running a magnetic field inversion software

Our goals are B maps spatial resolution better than 0.5" arcsec (3 times better than before, and comparable to HINODE results)

THEMIS new polarimetric analysis scheme

Full-Stokes analyzer (An4) **located at the F1 prime focus**, is designed to deliver **dual-beam polarimetry with beam exchange**.



Dual-beam (single Savart plate)

$$S_{+Q} = \begin{bmatrix} I \\ +Q \\ 0 \\ 0 \end{bmatrix} \quad S_{-Q} = \begin{bmatrix} I \\ -Q \\ 0 \\ 0 \end{bmatrix}$$

$$S_{+V} = \begin{bmatrix} I \\ +V \\ 0 \\ 0 \end{bmatrix} \quad S_{-V} = \begin{bmatrix} I \\ -V \\ 0 \\ 0 \end{bmatrix}$$

$$S_{+U} = \begin{bmatrix} I \\ +U \\ 0 \\ 0 \end{bmatrix} \quad S_{-U} = \begin{bmatrix} I \\ -U \\ 0 \\ 0 \end{bmatrix}$$

Beam exchange

Input

$$S = \begin{bmatrix} I \\ Q \\ U \\ V \end{bmatrix}$$

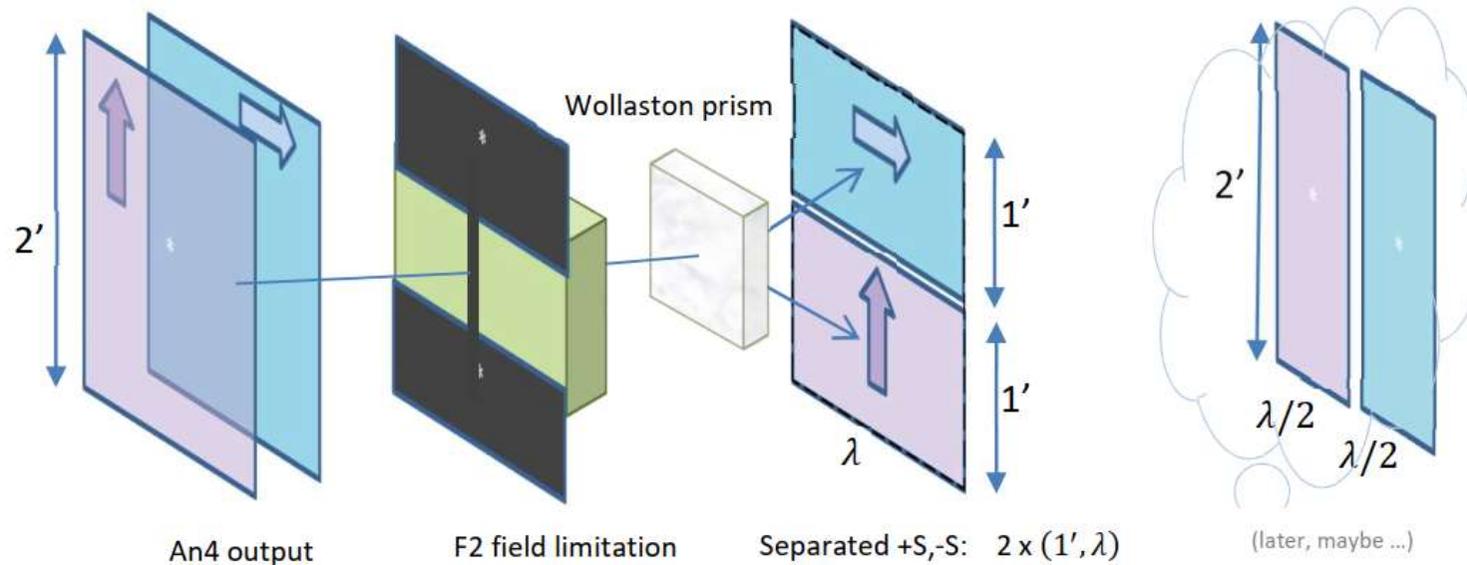
modulation (QW)

$$S_Q = \begin{bmatrix} I \\ Q \\ U \\ V \end{bmatrix} \quad S_V = \begin{bmatrix} I \\ V \\ Q \\ U \end{bmatrix} \quad S_U = \begin{bmatrix} I \\ U \\ V \\ Q \end{bmatrix}$$

We use a double Savart plate that generates the dual beam feature and **also superimpose both beams so that they behave as one**, differing only by their linear polarization state.

THEMIS new polarimetric analysis scheme

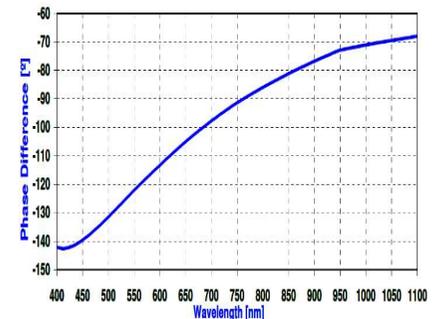
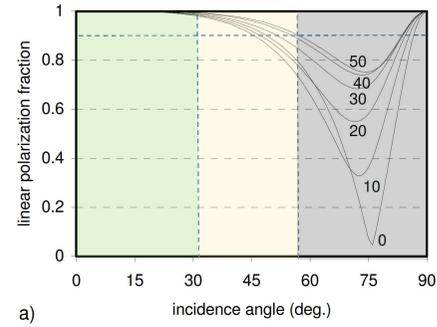
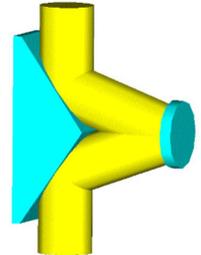
- Thanks to THEMIS “**polarization friendly**” new optical path (**geometry of the elevation axis, field rotator, coatings**), this output can travel through the telescope and reach the spectrograph cameras “minimally perturbed”
- Just in front of the spectral cameras**, a Wollaston prism splitter (one per camera) **separates the superimposed beam** into complementary Stokes components to form the spectral focal plane.



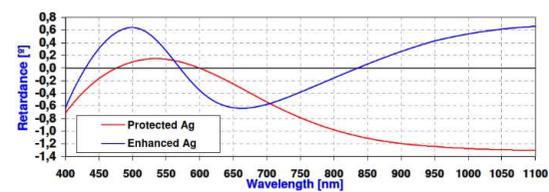
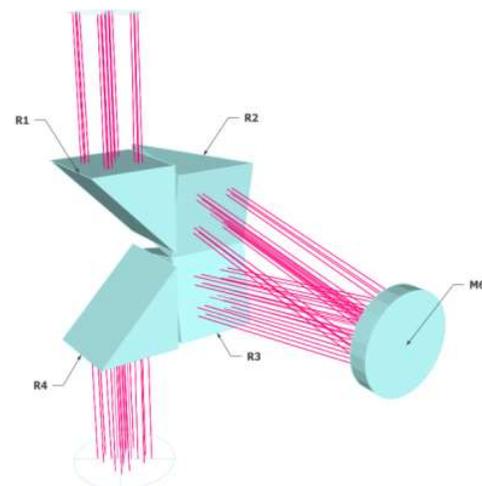
- 2 possible configs, the “easy one” is only used for now

Example of a “polarization friendly” field rotator

- **BAD (really bad ...)**
- Made of 3 mirrors with incidence = **55° / 20° / 55°**
- chromatic retardance is enormous (60°)

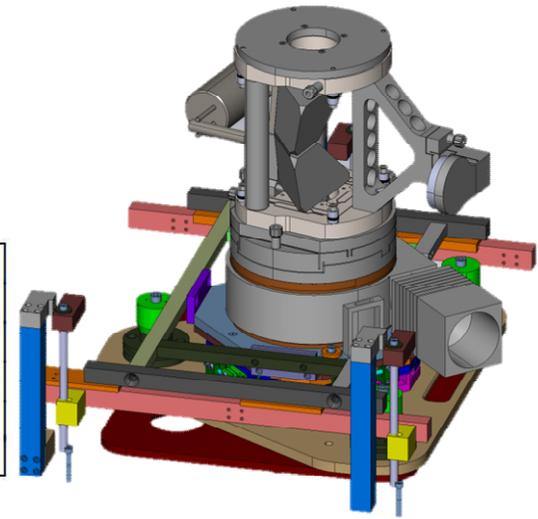


- **MUCH BETTER !!**
- Rotator, composed of 2 prisms and one mirror
incidence=45° / 45° / 10° / 45° / 45°
- **Retardance cancellation** of the crossed prisms@100°
(zero retardance for 90° only)
- **Coating required** on prisms hypotenuse

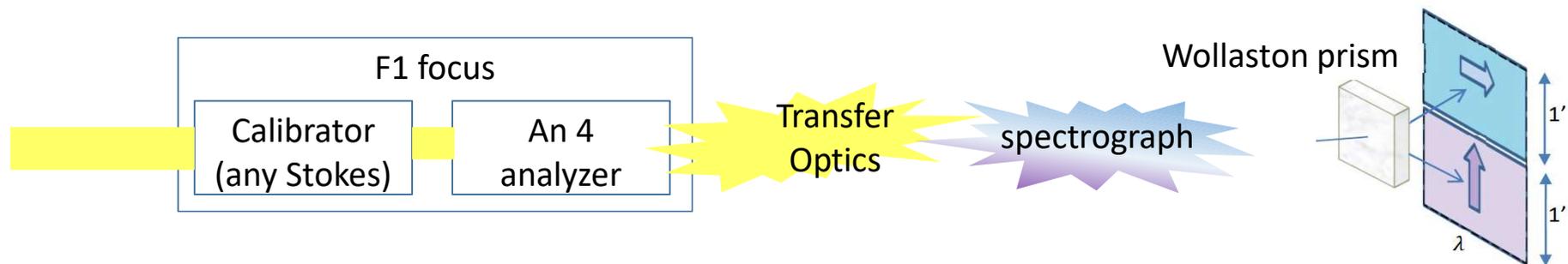


$$M_{Rot}(Ag) = \begin{bmatrix} 0.943 & 0.057 & 0 & 0 \\ 0.057 & 0.943 & 0 & 0 \\ 0 & 0 & 0.007 & -0.941 \\ 0 & 0 & 0.941 & 0.007 \end{bmatrix}$$

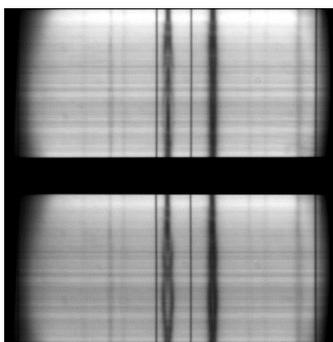
$$M_{ROT} = \begin{bmatrix} 0.9991 & 0.0009 & 0 & 0 \\ 0.0009 & 0.9859 & -0.0205 & 0.1605 \\ 0 & 0.0179 & 0.9989 & 0.0189 \\ 0 & -0.1607 & -0.0157 & 0.9859 \end{bmatrix}$$



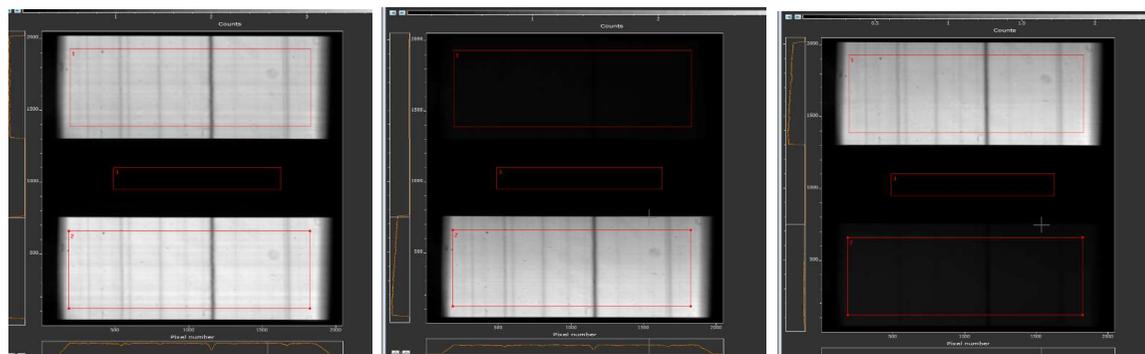
THEMIS polarimetric calibration



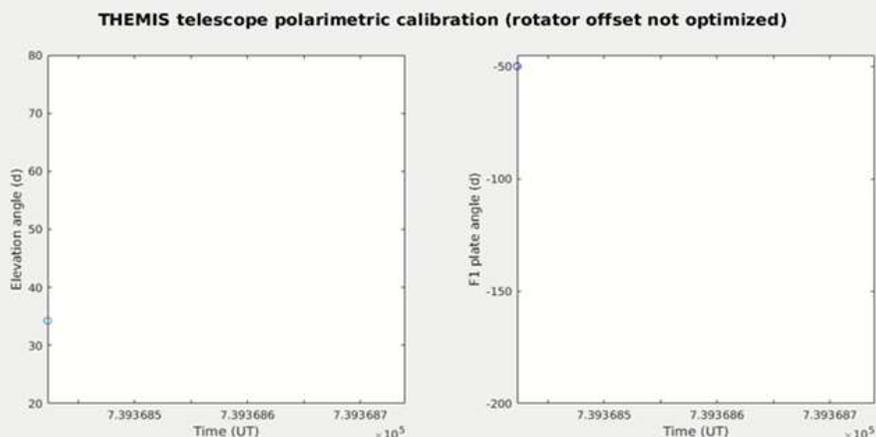
Analysis of natural solar spectrum



Analysis of calibrated solar spectrum

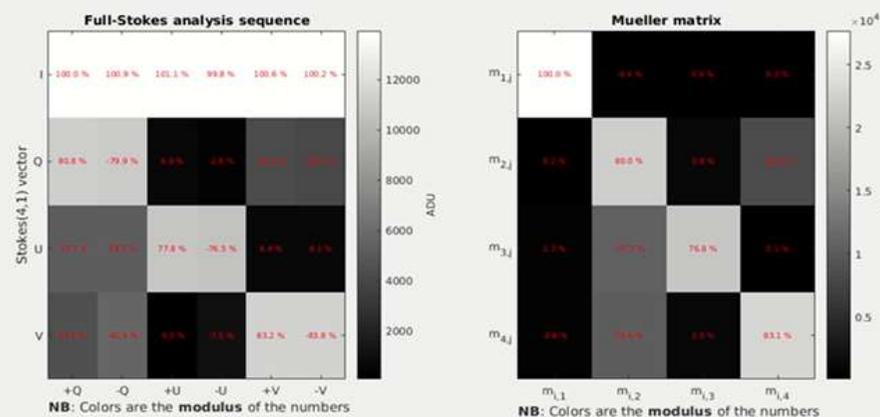


THEMIS Mueller matrix@~600nm



- $$M_{THEMIS} = \begin{pmatrix} 1.000 & -0.009 & -0.003 & 0.001 \\ -0.008 & 0.885 & 0.016 & -0.033 \\ 0.014 & -0.436 & 0.872 & 0.033 \\ -0.019 & 0.415 & 0.008 & 0.873 \end{pmatrix}$$

- Averaged over one full day
- Includes changing elevation axis and field derotation
- Quite constant along one day



Takeaway

- THEMIS is now a competitive XXIst century telescope, and the installation of the IBIS 2.0 spectroimager (end of 2025) will trigger an even larger europeanwide interest. THEMIS is a real challenger in the 1m to 1.5m class of solar telescopes and an a French and European asset.
- Several French teams have demonstrated their ability to successfully design, build, and operate on the sky new solar instrumentation using innovative concepts
- Several other French teams are leaders in the theoretical and numerical exploration of the solar magnetic field generation, surfacing, shaping of the corona configuration and driving the space weather events.
- We are clearly qualified in all the aspects of this business, and we are many (not “4 FTEs “ !!!)

... and nevertheless

THEMIS **IS NOT** the future of our community
because



the European Solar Telescope **IS THIS FUTURE**

Business continuity (... when shit happens...)

